Model for the height of the ionosphere maximum in main ionospheric trough zone

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Abstract. A model for spatial-temporal variations in the height of the ionospheric F2 region maximum in the main ionospheric trough (MIT) zone was developed via statistical analysis of data of external sounding of the subauroral ionosphere, obtained by the Intercosmos 19 satellite. This model is based on the IRI 90 model for $h_m F2$ and on a set of correcting factors depending on the level of magnetic activity and invariant latitude. Correction of the global ionosphere model IRI for the MIT zone allows us to reduce modeling error by half.

Introduction

Although many papers deal with investigations of the main ionospheric trough (MIT) [Ben'kova and Zikrach, 1983; Besprozvannaya et al., 1990; Dedeney et al., 1983; Moffett and Quegan, 1973; Rodger et al., 1992], relatively little attention is paid to variations in the height of the F2 region maximum $h_m F2$. Evidently, this lack of attention is related to the fact that capabilities of the network of ionosphere stations are limited because of the small number of stations in this region and the need for more frequent sounding when satellite data are employed. In the literature [Ben'kova and Zikrach, 1983; Besprozvannaya et al., 1990], experimental data for $h_m F2$ in the MIT region are given without general analysis of $h_m F2$ variations. Dedeney et al. [1983] gave a typical scheme for isolines of the electron density in the MIT region for the southern hemisphere; however, they did not reveal any peculiarities of the $h_m F2$ variations. The goal of this paper is to study spatial-temporal variations in $h_m F2$ on the basis of the external sounding data and also to construct an analytical model.

Experimental Data

The data from external sounding from aboard the Intercosmos 19 satellite for 1980–1982 (high solar activity, Wolf number W>120) were used for periods from winter to equinox in the northern hemisphere in the longitudinal sector 20 70°E. Ionograms were taken every 8 s, i.e., in intervals of about 40 km along the Earth's surface, which allowed the position and form of the MIT to be reliably determined. A total of 62 satellite flights

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through the MIT regions were studied, and each flight gave 20-45 ionograms, i.e., a total of 2000. We calculated altitude profiles of electron density and the height of the F2 region maximum via the procedure given by Indyukov and Serebryakova [1986], using the modified method of Jackson [1969].

Peculiaries of the Determination of $h_m F^2$ From External Sounding Ionograms

Usually the height of the F2 region maximum is determined from the external sounding ionograms as a true height at the last frequency which can be obtained from ionograms by the procedure given by Indyukov and Serebryakova [1986] and Jackson [1969]. The obtained estimates systematically exceed the h_mF2 values determined from ionograms of vertical sounding (VS) procedure [Gulyayeva, 1978], in which a parabolic extrapolation of a profile form near the F2 region maximum is adopted $[Ben'kova\ et\ al.,\ 1988]$. This excess increases with latitude (see Table 1), mainly for two reasons. The first is the existence of thick F2 layers with small or

Table 1. Parameter Δh_m versus latitude

VS station	Geographic	Mean
	latitude, deg	Δh_m , km
Yugorskiy Shar	70	+29
Arkhangel'sk	65	+30
Yakutsk	62	+18
Sofia	43	+18
Ashkhabad	38	+17
Tashkent	41	+17
Havana	23	+8

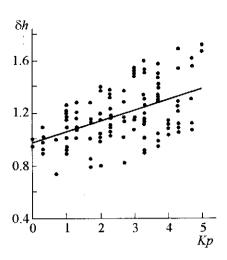


Figure 1. The $\delta h = (h_m F2)_{\rm exp}/\overline{h_m F2}$ versus Kp index within the range $-0.5^{\circ} \leq \Delta \Phi_s \leq 0.5^{\circ}$ in the vicinity of the MIT minimum. Solid circles represent experimental data; the solid line demonstrates the linear approximation through the least squares procedure.

zero vertical plasma frequency gradients in the vicinity of MIT. In this case a radio pulse from an onboard ionosonde is reflected from the topside of this region, and a pulse from a ground-based ionosonde is reflected from the bottom part of the region. As a result, the difference Δh_m is dominantly positive. The second reason for overestimation of $h_m F2$ is high diffusivity of external ionograms. Thus the procedure suggested by Indyukov and Screbryakova [1986] was supplemented by a parabolic extrapolation of the form of the external ionosphere profile, as was done by Gulyayeva [1978], and this allowed the systematic overestimation to be reduced. To extrapolate $h_m F2$, we used the following formulas:

$$h_m F2 = \overline{h_m F2} + (X1 - 1) \frac{h_{30} - \overline{h_m F2}}{X1 - X2}$$

$$X1 = 1 - \sqrt{1 - \frac{\overline{N_m F2}}{N_m F2}}$$

$$X2 = 1 - \sqrt{1 - \frac{N_{30}}{N_m F2}}$$

where h_mF2 is the height of the F2 region maximum without extrapolation and $\overline{N_mF2}$, N_mF2 and N_{30} , are the electron densities at the heights $\overline{h_mF2}$, h_mF2 , and $\overline{h_mF2+30}$ km, respectively.

Results of Statistical Analysis

Statistical analysis of $h_m F2$ variations, based on the external sounding data, and correlation between the experimental data and the data computed by the global

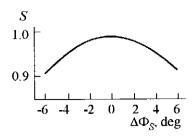


Figure 2. The function S, which demonstrates the latitudinal variations in relative values of δh , versus the MIT minimum position.

ionosphere model IRI 90 permitted the following conclusions.

- 1. The model gives systematically underestimated values for $h_m F2$ in the vicinity of the MIT minimum, the average error increasing with magnetic activity.
- 2. In the vicinity of the MIT minimum the ratio δh of the experimental values of $(h_m F2)_{\rm exp}$ to the calculated values $\overline{h_m F2}$ depends linearly on the magnetic activity level, characterized by a value of the Kp index. Figure 1 shows the δh distribution within the range $-0.5^{\circ} \leq \Delta \Phi_s \leq 0.5^{\circ}$ in the vicinity of the MIT minimum. Here $\Delta \Phi_s$ is the deviation of the invariant latitude of the sounding point from the MIT minimum position. The line demonstrates the linear approximation performed by the least squares procedure.
- 3. The latitudinal dependence of $h_m F2$ has a pronounced maximum near the MIT minimum position. Figure 2, which presents the ratio $S = \delta h/\overline{\delta h}$ dependence on $\Delta \Phi_s$, clearly demonstrates this finding. Here δh is determined at the current sounding point along the satellite orbit, and $\overline{\delta h}$ is associated with the MIT minimum position. This profile is in agreement with a typical diagram of a two-dimensional MIT section [Dedeney et al., 1983] as well as with theoretical results [Bryunelli and Namgaladze, 1988; Rodger et al., 1992; Watkins and Richards, 1979].

The initial heating of the thermosphere in the auroral zone, together with the increase of magnetic activity, resulting from the intensification of auroral electrojets and the absorption of energy from injected particles, forms a system of thermospheric winds directed from the heated region toward the equator. Under the action of a horizontal wind the plasma rises along the geomagnetic field lines, thereby increasing the height of the F2 region maximum in the MIT zone.

Analytical Model

Results of the statistical analysis form the basis for a model of spatial-temporal variations in the height of the F2 region maximum in the MIT zone. This model

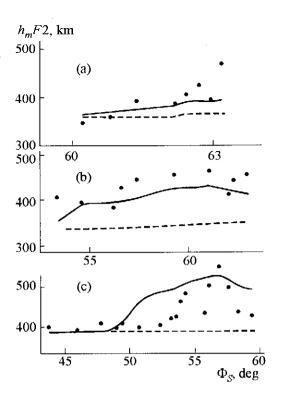


Figure 3. The h_mF2 variations versus an invariant geomagnetic latitude along the Intercosmos 19 orbit: (a) October 30, 1980, Kp = 1+; (b) March 18, 1981, Kp = 3+; and March 3, 1982, Kp = 5-. Circles represent experimental data; dashed lines show the data from the IRI 90 model; and solid lines show the data from the IRI model corrected for the MIT region.

uses the following analytical expressions:

$$h_m F2 = \overline{h_m F2} (\alpha + \beta K p) S(\Delta \Phi_s)$$
$$S(\Delta \Phi_s) = \alpha_1 (\Delta \Phi_s)^{\beta_1} \exp(\gamma_1 \Delta \Phi_s)$$

where $S(\Delta\Phi_s)$ is a function describing latitudinal variations in relative quantities δh within the range $-6^{\circ} \leq \Delta\Phi_s \leq 6^{\circ}$; $\alpha = 0.971$; $\beta = 0.081$; $\alpha_1 = 0.2896 \times 10^{-1}$; $\beta_1 = 1.783$; and $\gamma_1 = -0.9006 \times 10^{-1}$. This model can be easily included in the IRI 90 model, which was done by a smooth transition from the subauroral to the middle latitudes. Figure 3 presents samples of experimental measurements of $h_m F2$ variations along the satellite orbit and also their simulation using the original IRI 90 model, corrected for the MIT zone at various levels of magnetic activity. We can see that the correcting procedure decreases by half the relative error of simulation.

Conclusion

A model for spatial-temporal variations in the height of the F2 region maximum in the MIT zone was developed as a result of statistical analysis of data from external sounding of the subauroral ionosphere. The model is based on the IRI model for h_mF2 and on a set of correcting factors depending on the magnetic activity level. The corrected model for the MIT zone allows the simulation error to be reduced by half.

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